

β -DECAY

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✓

e^- ~~was~~ β^- DECAY WAS 1ST RADIATION DETECTOR
or discovered

e^+ OR β^+ DECAY WAS NOT OBSERVED UNTIL 1934
E.C. DISCOVERED 1938

ALL THREE PROCESSES LUMPED UNDER " β -DECAY"

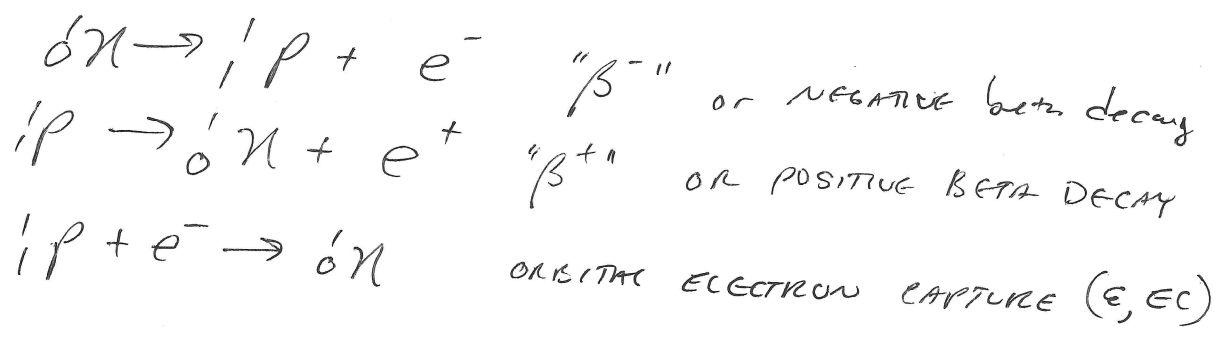
• BASIC β^- DECAY IS THE CONVERSION OF A PROTON INTO A n OR OF A $n \rightarrow$ INTO A PROTON.

• IN THE NUCLEI, β^- DECAY CHANGES BOTH Z & N BY ONE UNIT

$$Z \rightarrow Z \pm 1, N \rightarrow N \mp 1, A = Z + N = \text{CONSTANT}$$

• β^- DECAY PROCESS "CREATES" AN e^- FROM AVAILABLE DECAY ENERGY AT THE INSTANT OF DECAY (CONTRASTING THE PRE-FORMED α PARTICLE MODEL)

• BASIC DECAY PROCESSES ARE

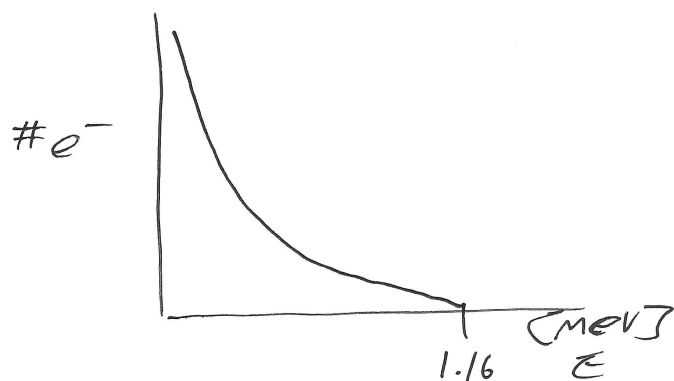


UNLIKE α -DECAY, ~~WHICH~~ WHOSE α -ENERGY ARE SHARPLY DEFINED, EQUAL TO MASS ENERGY DIFFERENCE BETWEEN INITIAL + FINAL STATES

β ~~particles~~ PARTICLES HAVE A CONTINUOUS DISTRIBUTION OF ENERGIES FROM 0 \rightarrow UPPER LIMIT KNOWN AS "END POINT" ENERGY

- IF β -DECAY WERE A 2-BODY PROCESS WE WOULD EXPECT ALL OF THE β -PARTICLE ENERGY TO BE UNIQUE + DISCRETE.

1.2. ^{210}Bi FROM MASS DIFFERENCE CALCULATION YOU'D EXPECT THE e^- TO HAVE K.E. OF 1.16 MeV YET THERE IS A CONTINUOUS SPECTRUM 0 \rightarrow UP TO 1.16 MeV



W. PAULI PROPOSED ANOTHER, NON-DETECTABLE, PARTICLE
ENRICO FERMI NAMED IT THE "NEUTRINO" 3/

- VERY PENETRATING RADIATION
- CHARGE CONSERVATION $\Rightarrow \bar{\nu}$ IS NEUTRAL
- CONS. OF ANG. MOMENTUM \Rightarrow SPIN $1/2$
(LIKE THE ELECTRON)

TWO TYPES OF NEUTRINOS: NEUTRINO + ANTI NEUTRINO

β^- IS ACCOMPANIED BY $\bar{\nu}$

$EC + \beta^+$ IS ACCOMPANIED BY ν

β -DECAY ENERGISTICS, CONSIDER DECAY OF
FREE NEUTRON $T_{1/2} \sim 15 \text{ MIN}$



DEFINE Q AS ~~THE~~ DIFFERENCE BETWEEN INITIAL &
FINAL NUCLEAR MASS ENERGIES

$$Q = (m_n - m_p - m_e - m_{\bar{\nu}})c^2 \quad + \text{ FOR DECAY @ REST}$$

$$Q = T_p + T_{e^-} + T_{\bar{\nu}}$$

IGNORE PROTON RECOIL ENERGY ($T_p \sim 0.3 \text{ keV}$) 4/

THE $e^- + \bar{\nu}$ WILL SHARE THE DECAY ENERGY.

ACCOUNTING FOR THE CONTINUOUS ENERGY SPECTRUM.

MAX e^- MIN $\bar{\nu}$ + VICE VERSA

IN β^- DECAY, MAX MEASURED e^- ENERGY

IS $0.782 \pm 0.03 \text{ MeV}$, ~~FROM WHICH~~ WE

CAN ALSO COMPUTE Q VALUE USING β^- , m_p , & e^- MASS VALUES

$$Q = m_n c^2 - m_p c^2 - m_e c^2 - m_{\bar{\nu}} c^2$$

$$= 939.573 \text{ MeV} - 938.280 \text{ MeV} - 0.511 \text{ MeV} - m_{\bar{\nu}} c^2$$

$$= 0.782 \text{ MeV} - m_{\bar{\nu}} c^2 \quad \nabla$$

FROM THIS WE CAN INFER $\bar{\nu}$ IS MASSLESS

IF WE ASSUME $\bar{\nu}, \nu$ IS MASSLESS, THEN IT MOVES AT THE SPEED OF LIGHT.

ONE CAN CONSIDER ITS TOTAL RELATIVISTIC $E_{\bar{\nu}}$ THE SAME AS ITS KINETIC ENERGY.

PART II

INTERACTION OF e^- W/ MATTER

6/

β -PARTICLES CAN EXCITE + IONIZE MATTER,

THEY CAN ALSO RADIATE BY "BREMSSTRAHLUNG"

BUT IS ONLY IMPORTANT HIGH ENERGIES.

~ 100 MEV RADIATIVE POWER IS $\sim 50\%$ IN H_2O

CONSIDER SEPARATELY COLLISIONAL STOPPING POWER

$$-\left(\frac{dE}{dx}\right)_{col}$$

+ RADIATIVE STOPPING POWER

$$-\left(dE/dx\right)_{rad}$$

β -PARTICLES CAN ALSO BE SCATTERED ELASTICALLY BY

ATOMIC e^- , A PROCESS THAT HAS A SIGNIFICANT

EFFECT ON β -PARTICLE PENETRATIONS

- CANNOT DISTINGUISH BETWEEN INCIDENT + STRUCK ELECTRON AFTER THE COLLISION, CONVENTION ~~IS~~ HAS ENERGY LOSS DEFINED AS e^- WITH LOWER ENERGY AFTER COLLISION IS "STRUCK" e^-

COLLISION STOPPING POWER

$$\left(-\frac{dE}{dx}\right)_{col}^{\pm} = \frac{4\pi k_0^2 e^4 n}{m c^2 \beta^2} \left[\ln \frac{m c^2 \tilde{\tau} \sqrt{\tilde{\tau}+2}}{\sqrt{2} I} + F^{\pm}(\beta) \right]$$

WHERE FOR ELECTRONS

$$F^{-}(\beta) = \frac{1-\beta^2}{2} \left[1 + \frac{\tilde{\tau}^2}{8} - (2\tilde{\tau}+1) \ln(\tilde{\tau}) \right]$$

+ FOR POSITRONS:

$$F^{+}(\beta) = \ln 2 - \frac{\beta^2}{24} \left[23 + \frac{14}{\tilde{\tau}+2} + \frac{10}{(\tilde{\tau}+2)^2} + \frac{4}{(\tilde{\tau}+2)^3} \right]$$

WHERE $\tilde{\tau} = \frac{T}{m c^2}$ + KE. OF β^{-} OR β^{+} PARTICLE IS EXPRESSED IN MULTIPLES OF e^{-} REST MASS ENERGY, $m c^2$.

$$\left(-\frac{dE}{dx}\right)_{col}^{\pm} = \frac{5.08 \times 10^{-31} n}{\beta^2} \left[\ln \frac{3.61 \times 10^5 \tilde{\tau} \sqrt{\tilde{\tau}+2}}{I_{ev}} + F^{\pm}(\beta) \right] \frac{MeV}{cm}$$

PUT INTO A GENERAL FORM

$$\left(-\frac{dE}{dx}\right)_{col}^{\pm} = \frac{5.08 \times 10^{-31} n}{\beta^2} \left[G^{\pm}(\beta) - \ln(I_{ev}) \right] \frac{MeV}{cm}$$

$$G^{\pm}(\beta) = \ln(3.61 \times 10^5 \tilde{\tau} \sqrt{\tilde{\tau}+2}) + F^{\pm}(\beta)$$

EXAMPLE

CALCULATE STOPPING POWER OF WATER

FOR 1-MeV e^-

8

SOLUTION: FIND $\left(-\frac{dE}{dx}\right)_{col}$ COMPUTE β^2 , $F(\beta)$, + $G(\beta)$ WE HAVE $N = 3.34 \times 10^{29} e^-/m^3$
 $\ln(I_{ev}) = 4.31$ FIND β^2 : $1 \text{ MeV} = 0.511 \text{ MeV} \left(\frac{1}{\sqrt{1-\beta^2}} - 1 \right)$

$$\beta^2 = 0.886$$

$$\tilde{z} = \frac{T}{Mc^2} = \frac{1}{0.511} = 1.96$$

$$\begin{aligned} \therefore F(\beta) &= \frac{1-0.886}{2} \left[1 + \frac{(1.96)^2}{8} - (2 \times 1.96 + 1) \ln(2) \right] \\ &= -0.11 \end{aligned}$$

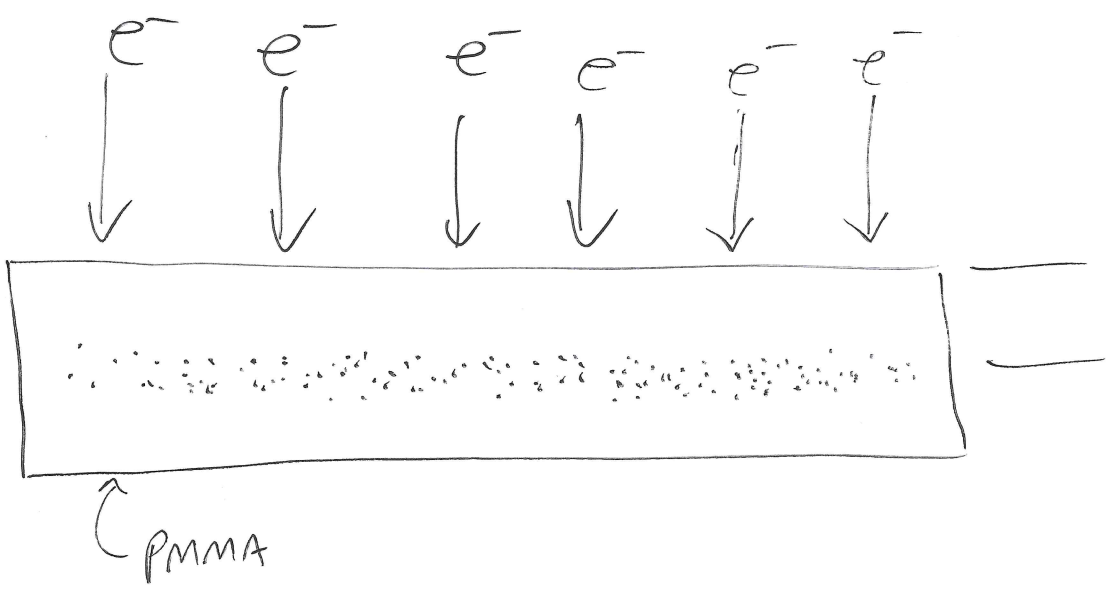
$$\therefore G(\beta) = \ln(3.34 \times 10^{29} \times 1.96 \sqrt{1.96+2}) - 0.110 = 14.0$$

PULL IT ALL TOGETHER:

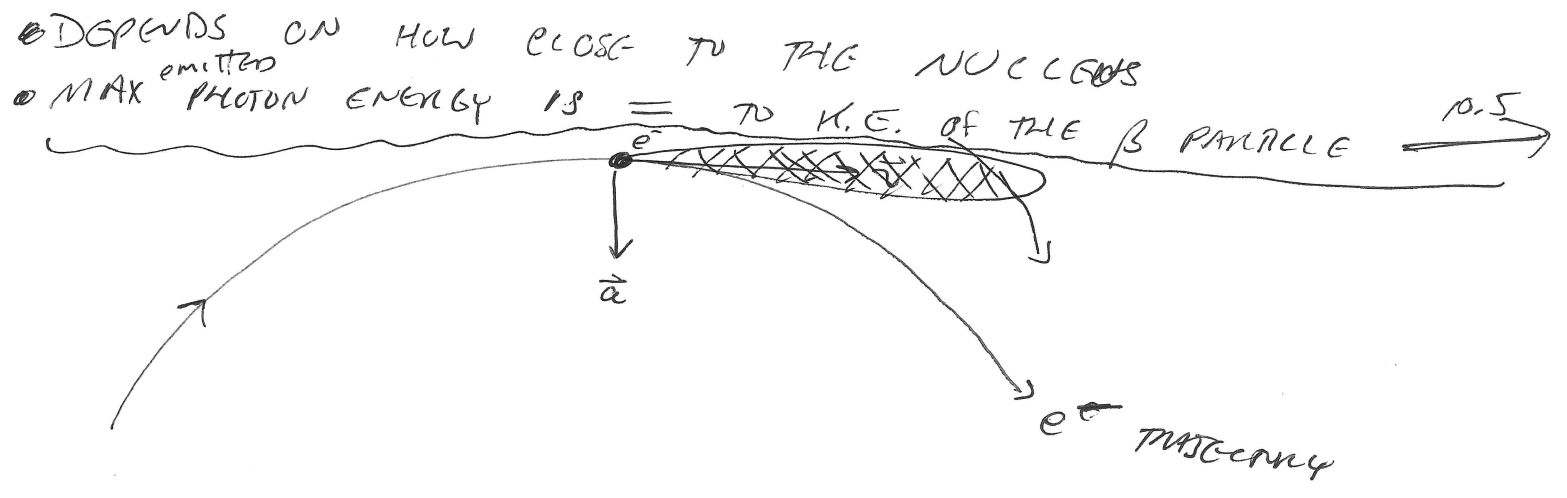
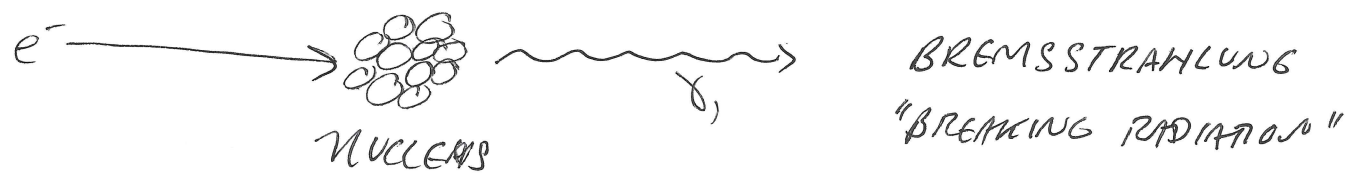
$$\begin{aligned} \left(-\frac{dE}{dx}\right)_{col} &= \frac{5.08 \times 10^{31} \times 3.34 \times 10^{29}}{0.886} [14 - 4.31] \\ &= 1.86 \text{ MeV/cm} \end{aligned}$$

LICHTENBERG FIGURE DEMONSTRATION + DISCUSSION

- DEPTH, THICKNESS



RADIATING STOPPING POWER



EXAMPLES OF LIGHT SOURCES

CALCULATING RADIATIVE STOPPING POWER

$\left(-\frac{dE}{dx}\right)_{\text{rad}}$ IS FROM NUMEROUS

PROCESSES. THE PROCESS IS VERY DIFFERENT FROM IONIZATION & EXCITATION

• THE EFFICIENCY OF BREMSTRAHLUNG IN DIFFERENT ELEMENTS VARIES WITH Z^2 .

• THUS "CONVERSION" TO PHOTON IS MUCH GREATER THAN WITH HIGH Z -MATERIAL, SUCH AS LEAD

• APPROXIMATE RATIO OF RADIATIVE & COLLISIONAL STOPPING POWERS FOR AN ELECTRON TOTAL ENERGY E , IN MeV:

$$\frac{\left(-\frac{dE}{dx}\right)_{\text{rad}}}{\left(-\frac{dE}{dx}\right)_{\text{col}}} \approx \frac{ZE}{800}$$

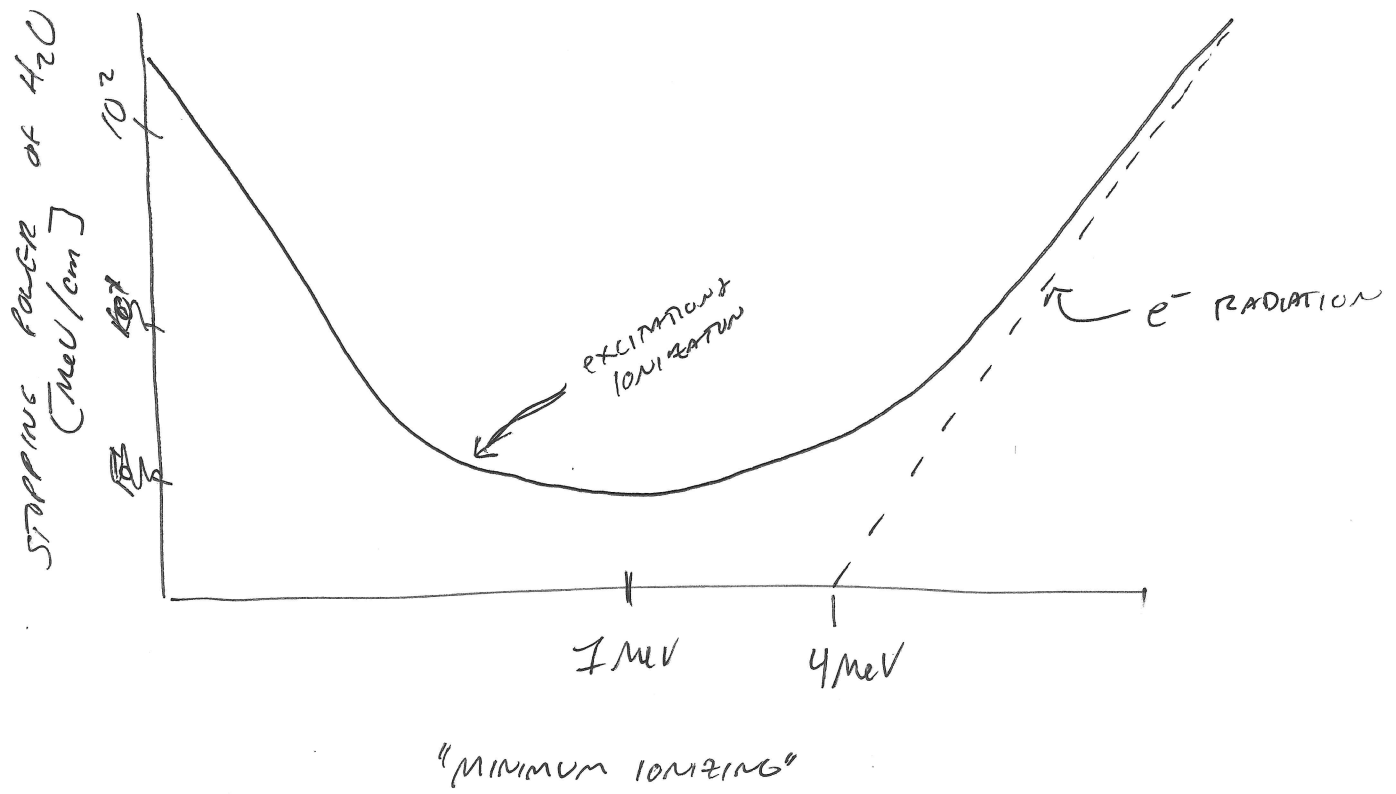
FOR $Z = 82$:

$$\frac{82E}{800} \approx 1 \text{ MeV} \rightarrow E \approx 9.8 \text{ MeV}$$

$$\text{K.E. } T = E - mc^2 \approx 9.3 \text{ MeV}$$

OXYGEN $Z = 8$, $E \approx 100 \text{ MeV}$

IN WATER:



• AN ESTIMATE OF RADIATION YIELD ~~CONCEPTS~~

INITIAL e^- KE IS T (MeV) STOPPING IN Z

THEY RADIATION YIELD IS

$$Y \approx \frac{6 \times 10^{-4} Z T}{1 + 6 \times 10^{-4} Z T}$$

EXAMPLE: 2 MeV β SHIELDED BY AL. & Pb

Al: $ZT = 13 \times 2 = 26$ $Y \approx 0.016 / 1.016 = 0.016 \rightarrow 1.6\% \text{ IN RAD.}$

Pb: $ZT = 82 \times 2 = 164$ $Y = 0.090 \rightarrow 9\%$

RANGE SIMILAR TO ~~LEAVE~~ PARTICLES, e^- RANGE

"R" = RANGES ARE EXPRESSED IN

$$[g/cm^2]$$

FOR $0.01 \leq T \leq MeV$

$$R = 0.412 T^{1.27 - 0.0954 \ln(T)}$$

FOR $T > 2.5 MeV$

$$R = 0.530 T - 0.106$$

$$d = \frac{R}{\rho} = [cm]$$

EXAMPLE: 2.2 MeV e^- IN LUCITE:

$$R = (0.412)(2.2)^{1.27 - 0.0954 \ln(2.2)} = 1.06 g/cm^2$$

$$d = \frac{R}{\rho} = \frac{1.06 g/cm^2}{1.19 g/cm^3} = 0.891 cm$$

OPTIONAL — LAB —

DISCUSS LAB